## Exercise 21

Assume that the earth is a solid sphere of uniform density with mass $M$ and radius $R=3960 \mathrm{mi}$. For a particle of mass $m$ within the earth at a distance $r$ from the earth's center, the gravitational force attracting the particle to the center is

$$
F_{r}=\frac{-G M_{r} m}{r^{2}}
$$

where $G$ is the gravitational constant and $M_{r}$ is the mass of the earth within the sphere of radius $r$.
(a) Show that $F_{r}=\frac{-G M m}{R^{3}} r$.
(b) Suppose a hole is drilled through the earth along a diameter. Show that if a particle of mass $m$ is dropped from rest at the surface, into the hole, then the distance $y=y(t)$ of the particle from the center of the earth at time $t$ is given by

$$
y^{\prime \prime}(t)=-k^{2} y(t)
$$

where $k^{2}=G M / R^{3}=g / R$.
(c) Conclude from part (b) that the particle undergoes simple harmonic motion. Find the period $T$.
(d) With what speed does the particle pass through the center of the earth?

## Solution

The mass of the earth is its density times its volume.

$$
M=\rho\left(\frac{4}{3} R^{3}\right)
$$

As a result, the earth's density is

$$
\rho=\frac{M}{\frac{4}{3} R^{3}} .
$$

The mass of earth within a sphere of radius $r$ is

$$
\begin{aligned}
M_{r} & =\rho\left(\frac{4}{3} r^{3}\right) \\
& =\left(\frac{M}{\frac{4}{3} R^{3}}\right)\left(\frac{4}{3} r^{3}\right) \\
& =M\left(\frac{r^{3}}{R^{3}}\right) .
\end{aligned}
$$

Consequently, the gravitational force attracting the particle to the center of the earth is

$$
\begin{aligned}
F_{r} & =\frac{-G M_{r} m}{r^{2}} \\
& =\frac{-G m}{r^{2}}\left[M\left(\frac{r^{3}}{R^{3}}\right)\right] \\
& =\frac{-G M m}{R^{3}} r .
\end{aligned}
$$

Apply Newton's second law to get the equation of motion for a mass that oscillates through the center of the earth.

$$
\sum F=m a
$$

The only force acting on the mass is the gravitational force.

$$
F_{r}=m a
$$

Substitute the formula for $F_{r}$ and use the fact that acceleration is the second derivative of position.

$$
\frac{-G M m}{R^{3}} y=m \frac{d^{2} y}{d t^{2}}
$$

Divide both sides by $m$.

$$
\frac{d^{2} y}{d t^{2}}=-\frac{G M}{R^{3}} y
$$

Since the mass is dropped from rest at the earth's surface, the initial conditions associated with this ODE are $y(0)=R$ and $y^{\prime}(0)=0$. Set $k^{2}=G M / R^{3}$.

$$
y^{\prime \prime}=-k^{2} y
$$

This is a linear homogeneous ODE with constant coefficients, so it has solutions of the form $y=e^{p t}$.

$$
y=e^{p t} \quad \rightarrow \quad y^{\prime}=p e^{p t} \quad \rightarrow \quad y^{\prime \prime}=p^{2} e^{p t}
$$

Substitute these formulas into the ODE.

$$
p^{2} e^{p t}=-k^{2}\left(e^{p t}\right)
$$

Divide both sides by $e^{p t}$.

$$
p^{2}=-k^{2}
$$

Solve for $p$.

$$
p=\{-i k, i k\}
$$

Two solutions to the ODE are $y=e^{-i k t}$ and $y=e^{i k t}$. According to the principle of superposition, the general solution is a linear combination of these two.

$$
\begin{aligned}
y(t) & =C_{1} e^{-i k t}+C_{2} e^{i k t} \\
& =C_{1}(\cos k t-i \sin k t)+C_{2}(\cos k t+i \sin k t) \\
& =\left(C_{1}+C_{2}\right) \cos k t+\left(-i C_{1}+i C_{2}\right) \sin k t \\
& =C_{3} \cos k t+C_{4} \sin k t
\end{aligned}
$$

Differentiate it with respect to $t$.

$$
y^{\prime}(t)=-C_{3} k \sin k t+C_{4} k \cos k t
$$

Apply the initial conditions to determine $C_{3}$ and $C_{4}$.

$$
\begin{aligned}
y(0) & =C_{3}=R \\
y^{\prime}(0) & =C_{4} k=0
\end{aligned}
$$

Solving this system yields $C_{3}=R$ and $C_{4}=0$. Therefore,

$$
y(t)=R \cos k t .
$$

The period of the particle is

$$
T=\frac{2 \pi}{k}=\frac{2 \pi}{\sqrt{\frac{G M}{R^{3}}}}=2 \pi \sqrt{\frac{R^{3}}{G M}} .
$$

Notice that the particle reaches the center of the earth when $y(t)=0$, or when $k t=\pi / 2$. To find the speed that the particle has as it goes through the center of the earth, evaluate $\left|y^{\prime}\left(\frac{\pi}{2 k}\right)\right|$.

$$
y^{\prime}(t)=-R k \sin k t \quad \Rightarrow \quad\left|y^{\prime}\left(\frac{\pi}{2 k}\right)\right|=R k=R \sqrt{\frac{G M}{R^{3}}}=\sqrt{\frac{G M}{R}}
$$

The constants have the following numerical values.

$$
\begin{aligned}
M & =5.9736 \times 10^{24} \mathrm{~kg} \\
R & =6.378136 \times 10^{6} \mathrm{~m} \\
G & =6.67384 \times 10^{-11} \frac{\mathrm{~N} \cdot \mathrm{~m}^{2}}{\mathrm{~kg}^{2}}
\end{aligned}
$$

That means the period and speed are

$$
\begin{aligned}
& T \approx 5068.91 \phi \times \frac{1 \mathrm{~min}}{60 \phi} \approx 84.5 \text { minutes }
\end{aligned}
$$

